#### UNIVERSITY OF CAMBRIDGE INTERNATIONAL EXAMINATIONS

GCE Advanced Subsidiary Level and GCE Advanced Level

# MARK SCHEME for the October/November 2010 question paper for the guidance of teachers

## 9702 PHYSICS

9702/42

Paper 4 (A2 Structured Questions), maximum raw mark 100

This mark scheme is published as an aid to teachers and candidates, to indicate the requirements of the examination. It shows the basis on which Examiners were instructed to award marks. It does not indicate the details of the discussions that took place at an Examiners' meeting before marking began, which would have considered the acceptability of alternative answers.

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### **Section A**

1	(a) for	ce per unit mass	(ratio idea essential)	B1	[1]

(b) graph: correct curvature M1 from 
$$(R,1.0\,g_{\rm S})$$
 & at least one other correct point A1 [2]

(c) (i) fields of Earth and Moon are in opposite directions

either resultant field found by subtraction of the field strength

or any other sensible comment

so there is a point where it is zero

(allow 
$$F_E = -F_M$$
 for 2 marks)

(ii) 
$$GM_E/x^2 = GM_M/(D-x)^2$$
 C1  
 $(6.0 \times 10^{24})/(7.4 \times 10^{22}) = x^2/(60R_E-x)^2$  C1  
 $x = 54R_E$  A1 [3]

(iii) graph: 
$$g = 0$$
 at least  $\frac{2}{3}$  distance to Moon
$$g_{\rm E} \text{ and } g_{\rm M} \text{ in opposite directions}$$

$$\text{correct curvature (by eye) and } g_{\rm E} > g_{\rm M} \text{ at surface}$$

$$\text{A1} \quad [3]$$

- 2 (a) (i) no forces (of attraction or repulsion) between atoms / molecules / particles B1 [1]
  - (ii) sum of kinetic and potential energy of atoms / molecules M1 due to random motion A1 [2]
  - (iii) (random) kinetic energy increases with temperature no potential energy (so increase in temperature increases internal energy)

    A1 [2]
  - (b) (i) zero A1 [1]

(ii) work done = 
$$p\Delta V$$
 C1  
=  $4.0 \times 10^5 \times 6 \times 10^{-4}$   
= 240 J (ignore any sign) A1 [2]

(iii)

change	work done / J	heating / J	increase in internal energy / J
$\begin{array}{c} P \rightarrow Q \\ Q \rightarrow R \\ R \rightarrow P \end{array}$	<b>+240</b>	-600	-360
	0	+720	+720
	<b>-840</b>	+480	-360

(correct signs essential)
(each horizontal line correct, 1 mark – max 3)

B3 [3]

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3	(a)	(i)	resor	nance		B1	[1]
	(	(ii)	ampl	A1	[1]		
	(b)	(i)	a =	$(-)\omega^2 x$ and $\omega = 2\pi f$ $4\pi^2 \times 4.6^2 \times 16 \times 10^{-3}$ $13.4 \mathrm{m  s^{-2}}$		C1 C1 A1	[3]
	(	(ii)	F =	<i>ma</i> 150 × 10 <sup>−3</sup> × 13.4		C1	
				2.0N		A1	[2]
				ys 'below' given line and never zero t 4.6 Hz (or slightly less) and flatter		M1 A1	[2]
4	(a)	cha	rge / p	potential (difference) (ratio must be clear)		B1	[1]
	(b)	(i)	V = (	Q / 4πε <sub>0</sub> r		B1	[1]
	(	(ii)	<i>C</i> = ( so <i>C</i>	$Q/V = 4\pi \varepsilon_0 r$ and $4\pi \varepsilon_0$ is constant $\propto r$		M1 A0	[1]
	(c)	(i)	r = (6	$/4\pi\epsilon_0 r$ $6.8 \times 10^{-12}$ ) / $(4\pi \times 8.85 \times 10^{-12})$ $\times 10^{-2}$ m		C1 C1 A1	[3]
	(	(ii)		$CV = 6.8 \times 10^{-12} \times 220$ = 1.5 × 10 <sup>-9</sup> C		A1	[1]
	(d)	(i)	V = 0 = 83	$Q/C = (1.5 \times 10^{-9}) / (18 \times 10^{-12})$		A1	[1]
	(	(ii)	eithe	r energy = $\frac{1}{2}CV^2$ $\Delta E = \frac{1}{2} \times 6.8 \times 10^{-12} \times 220^2 - \frac{1}{2} \times 18 \times 10^{-12} \times 83^2$		C1 C1	
			or	= $1.65 \times 10^{-7} - 6.2 \times 10^{-8}$ = $1.03 \times 10^{-7}$ J energy = $\frac{1}{2}$ QV $\Delta E = \frac{1}{2} \times 1.5 \times 10^{-9} \times 220 - \frac{1}{2} \times 1.5 \times 10^{-9} \times 83$ = $1.03 \times 10^{-7}$ J		A1 (C1) (C1) (A1)	[3]

Mark Scheme: Teachers' version

Syllabus

Paper

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5	(a)	) field into		(the plane of) the paper		B1	[1]
	(b)		? / r = = (20	e to magnetic field <u>provides</u> the centripetal force $Bqv$ $0 \times 1.66 \times 10^{-27} \times 1.40 \times 10^{5}$ ) / (1.6 × 10 <sup>-19</sup> × 6.4 × 10 <sup>-14</sup> )	<sup>2</sup> )	B1 C1 B1 A0	[3]
	(c)	(i)	<u>sem</u>	icircle with diameter greater than 12.8 cm		B1	[1]
		(ii)	new	flux density = $\frac{22}{20} \times 0.454$		C1	
				$B = 0.499 \mathrm{T}$		A1	[2]
6	(a)	(i)	e.g.	prevent flux losses / improve flux linkage		B1	[1]
		(ii)	e.m.	in core is changing f. / current (induced) <u>in core</u> ced current in core causes heating		B1 B1 B1	[3]
	(b)	(i)		value of the direct current producing same (mean) pov resistor	ver / heating	M1 A1	[2]
		(ii)	•	er in primary = power in secondary $_{\rm o}$ = $V_{\rm S}I_{\rm S}$		M1 A1	[2]
7	(a)	(i)	e.g.	electron / particle diffraction		B1	[1]
		(ii)	e.g.	photoelectric effect		B1	[1]
	(b)	(i)		40		A1	[1]
		(ii)	$\lambda = I$	nge in energy = $4.57 \times 10^{-19}$ J hc / E $63 \times 10^{-34} \times 3.0 \times 10^{8}$ / ( $4.57 \times 10^{-19}$ )		C1	
			= 4.4	$4 \times 10^{-7} \mathrm{m}$		A1	[2]
8	(a)	-	_	of a heavy nucleus (not atom/nuclide) (lighter) nuclei of approximately same mass		M1 A1	[2]
	(b)	<sup>1</sup> <sub>0</sub> n <sup>4</sup> <sub>2</sub> He <sup>7</sup> <sub>3</sub> Li	Э	(allow ${}^4_2\alpha$ )		M2 A1	[3]
	(c)	rang lose	ge of kine	particles have kinetic energy particles in the control rods is short / particles stopped tic energy in rods nergy of particles converted to thermal energy	in rods /	B1 B1 B1	[3]

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## **Section B**

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		GCE AS	S/A LEVEL – October/November 2010	9702	42	
	(ii) use an amplifier coupled to the microphone (repeater amplifiers scores no mark)				M1 A1	[2]
12 (a	satellite i signal ar at a diffe different e.g. of fro	satellite receives greatly attenuated signal (1 signal amplified and transmitted <u>back to Earth</u> at a different (carrier) frequency different frequencies prevent swamping of uplink signal (1		(1) (1) (1) (1)	B1 B1 B2	[4]
(k	advantag	e.g.	because orbits are much lower whole Earth may be covered in several orbits / with network		M1 A1 (M1) (A1)	
	disadvar	naye. e.g.	<ul><li>either must be tracked</li><li>or limited use in any one orbit</li><li>more satellites required for continuous of</li></ul>	peration	M1 A1	[4]